

# Summary of OPTI 505R

## Chapter 9~14

by Dae Wook Kim

### I. Direct Phase Measurement Interferometry

#### 1.1 Advantages

- High measurement accuracy (>1/1000 fringe, fringe following only 1/10 fringe)
- Rapid measurement
- Good results with low contrast fringes
- Results independent of intensity variations across pupil
- Phase obtained at fixed grid of points
- Easy to use with large solid-state detector arrays

#### 1.2 Four Step Method

$$I(x,y) = I_{dc} + I_{ac} \cos[\underbrace{\phi(x,y)}_{\text{measured object phase}} + \underbrace{\phi(t)}_{\text{phase shift}}]$$

$I_1(x,y) = I_c + I_s \cos[\phi(x,y)]$	$\phi(t) = 0$ ( $0^\circ$ )
$I_2(x,y) = I_c - I_s \sin[\phi(x,y)]$	$= \pi/2$ ( $90^\circ$ )
$I_3(x,y) = I_c - I_s \cos[\phi(x,y)]$	$= \pi$ ( $180^\circ$ )
$I_4(x,y) = I_c + I_s \sin[\phi(x,y)]$	$= 3\pi/2$ ( $270^\circ$ )

$$\tan[\phi(x,y)] = \frac{I_1(x,y) - I_3(x,y)}{I_1(x,y) - I_2(x,y)}$$

Then, the height can be calculated as

$$\text{Height}(x,y) = \frac{\lambda}{4\pi} \phi(x,y)$$

#### 1.3 Phase Shifting Methods

Diffraction grating, Bragg Cell, Moving mirror, Rotating half wave plate, Rotating Polarizer Phase Shifter

#### 1.4 Phase Ambiguities

If we know sign of the Sin and sign of the Cosine the Arc Tangent is calculated modulo  $2\pi$ .

Removing Phase Ambiguities by

When phase jumps by  $> \pi$

Add or subtract  $N2\pi$

Adjust so  $< \pi$

#### 1.5 Way to fix the vibration problem

- Reduce vibration
- Take data fast
- Measure vibration and introduce vibration 180 degrees out of phase to cancel vibration
- Take all frames at once

#### 1.6 Vertical Scanning Interferometry

- Eliminates ambiguities in heights present with monochromatic interferometry
- Best fringe contrast corresponds to zero optical path difference
- Best focus corresponds to zero optical path difference

## II. Basic Diffraction Theories

### 2.1 Huygens Wavelet

Huygens wavelet can be drive using the Dirichlet B.C.m which is shown in the table.

Boundary Conditions	$G(P_0; P_1)$	Across $\Sigma$	In Shadow
Kirchhoff <sup>1</sup> (Cauchy)	$\frac{e^{jk_0 r_0}}{4\pi r_0}$	$U$ and $\frac{\partial U}{\partial n}$ same as without screen	$U=0$ $\frac{\partial U}{\partial n} = 0$
Neumann	$\frac{e^{jk_0 r_0}}{4\pi r_0} - \frac{e^{jk_0 r_0}}{4\pi r_0}$	$\frac{\partial U}{\partial n}$ same as without screen	$\frac{\partial U}{\partial n} = 0$
Dirichlet	$\frac{e^{jk_0 r_0}}{4\pi r_0} + \frac{e^{jk_0 r_0}}{4\pi r_0}$	$U$ same as without screen	$U=0$

The resulting Huygens wavelet is

$$h(P_0; P_1) = \frac{\partial}{\partial z_1} \frac{e^{jk_0 r_0}}{2\pi r_0} = \frac{\sqrt{1 + (kr_0)^2}}{2\pi r_0^2} \gamma_z e^{j[kr_0 - \tan^{-1}(kr_0)]}$$

where  $\gamma_z = \frac{z_0}{r_0}$

For  $r_0 \gg \lambda$  and  $\gamma_z \approx 1$ , it has Spherical wavefront

### 2.2 Fresnel wavelet

With approximation  $r_0 \gg \lambda$  and  $\frac{1}{r_0^2} = \frac{1}{z_0^2}$   $\exp(-j\phi)$  goes to the limiting value of  $-j$ .

Then, we get the Fresnel wavelet

$$h(P_0; P_1) = \frac{-j e^{jk_0 z_0}}{\lambda z_0} \exp\left\{ \frac{jk}{2z_0} [(x_0 - x_1)^2 + (y_0 - y_1)^2] \right\}$$

, which has Parabolic wavefront

The Fresnel diffraction formula is

$$U(x_0, y_0) = \frac{-j e^{jk_0 z_0}}{\lambda z_0} \exp\left[ \frac{jk}{2z_0} (x_0^2 + y_0^2) \right] \mathbf{F}_y \mathbf{F}_x \left\{ U_{z=0}(x_1, y_1) \exp\left[ \frac{jk}{2z_0} (x_1^2 + y_1^2) \right] \right\}$$

where  $\xi = x_0/\lambda z_0$ ,  $\eta = y_0/\lambda z_0$

This approximation is valid when

$$z_0 \gg \sqrt{\frac{\pi}{4} \lambda \left[ \left( \frac{x_0 - x_1}{\lambda} \right)^2 + \left( \frac{y_0 - y_1}{\lambda} \right)^2 \right]_{\max}^{2/3}}$$

However, the Fresnel approximation is valid over a larger range than the above condition, due to the "Principal of Stationary Phase."

### 2.3 Fraunhofer wavelet

With approximation  $z_0 \gg \frac{k}{2} (x_1^2 + y_1^2)_{\max}$

Then, we get the Fraunhofer wavelet

$$h(P_0; P_1) = \frac{-j e^{jk_0 z_0}}{\lambda z_0} e^{j \frac{k}{2z_0} (x_0^2 + y_0^2)} \exp\left[ \frac{2\pi j}{\lambda z_0} (x_0 x_1 + y_0 y_1) \right]$$

, which has Planar wavefront.

The Fraunhofer diffraction formula is

$$U(x_0, y_0) = \frac{-j e^{jk_0 z_0}}{\lambda z_0} e^{j \frac{k}{2z_0} (x_0^2 + y_0^2)} \mathbf{F}_y \mathbf{F}_x \left[ U_{z=0}(x_1, y_1) \right]$$

In fact, at the observation plane, direction cosines

good

$$\begin{aligned} \alpha &= \lambda \xi \\ \beta &= \lambda \eta \\ \gamma &= \sqrt{1 - \alpha^2 - \beta^2} \\ &= \sqrt{1 - (\lambda \xi)^2 - (\lambda \eta)^2} \end{aligned}$$

determine the angular orientation of each plane wave.

### 2.4 Diffraction using Transfer function

Using the "exact" Huygens wavelet", the transfer function of free space is

$$\mathbf{F}_\eta \mathbf{F}_\xi [h(P_0; P_1)] = e^{jk_0 z_0} = H(\gamma; z_0)$$

The diffraction formula is

$$\mathbf{F}_\eta \mathbf{F}_\xi [U(P_0)] = \mathbf{F}_\eta \mathbf{F}_\xi [U_{z=0}(P_1)] e^{jk_0 z_0}$$

Also, we can define "spatial frequency spectrum" (or "Angular spectrum of plane waves")

$$\mathbf{F}_\eta \mathbf{F}_\xi [U_{z=0}(P_1)] = \iint U_{z=0}(x_1, y_1) e^{-j2\pi(\xi x_1 + \eta y_1)} dx_1 dy_1 = A_{z=0}(\xi, \eta)$$

For an example, given

$$U_{z=0}(x_1) = \cos[2\pi(\alpha_0/\lambda)x_1] = \cos(2\pi\xi_0 x_1)$$

The "Angular spectrum of plane waves" is

$$\begin{aligned} A_{z=0}(\xi, \eta) &= \iint \cos(2\pi\xi_0 x_1) e^{-j2\pi(\xi x_1 + \eta y_1)} dx_1 dy_1 \\ &= \frac{1}{2} [\delta(\xi - \xi_0, \eta) + \delta(\xi + \xi_0, \eta)] \end{aligned}$$

The field at the observation plane is

$$\begin{aligned} U(x_0, y_0; z_0) &= \mathbf{F}_y^{-1} \mathbf{F}_x^{-1} [A_{z=0}(\xi, \eta) e^{jk_0 z_0}] \\ &= \mathbf{F}_y^{-1} \mathbf{F}_x^{-1} \left\{ \frac{1}{2} [\delta(\xi - \xi_0, \eta) + \delta(\xi + \xi_0, \eta)] e^{jk_0 z_0} \right\} \\ &= \frac{1}{2} \mathbf{F}_y^{-1} \mathbf{F}_x^{-1} \left[ e^{jk_0 \sqrt{1 - (\lambda \xi)^2 - (\lambda \eta)^2}} \delta(\xi - \xi_0, \eta) \right. \\ &\quad \left. + e^{jk_0 \sqrt{1 - (\lambda \xi)^2 - (\lambda \eta)^2}} \delta(\xi + \xi_0, \eta) \right] \end{aligned}$$

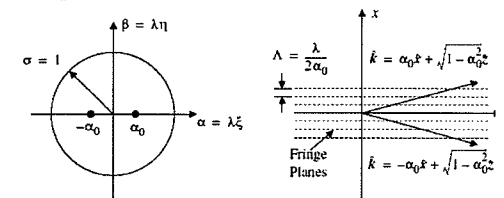
$$= \frac{1}{2} \left[ e^{jk_0 \sqrt{1 - (\lambda \xi_0)^2}} e^{j2\pi \xi_0 x_0} + e^{jk_0 \sqrt{1 - (\lambda \xi_0)^2}} e^{-j2\pi \xi_0 x_0} \right]$$

$$= \frac{1}{2} \left[ e^{j2\pi \xi_0 x_0} + e^{-j2\pi \xi_0 x_0} \right] e^{jk_0 \sqrt{1 - (\lambda \xi_0)^2}}$$

$$= \cos(2\pi \xi_0 x_0) e^{jk_0 \sqrt{1 - (\lambda \xi_0)^2}}$$

$$= \cos(k \alpha_0 x_0) e^{jk_0 \sqrt{1 - \alpha_0^2}}$$

We can think this result as combination of two plane waves



$$\text{, where } \sigma = (\alpha^2 + \beta^2)^{1/2}$$

$|\sigma| \leq 1$  : Propagating component

$|\sigma| > 1$  : evanescent components

### III. Important Diffraction Topics

#### 3.1 Talbot Effect

If the period of object is  $\gg$  wavelength

The maximum extent of the object's angular spectrum is limited to small angles:  $\sigma \ll 1$

For the Fraunhofer region, the free space transfer function H becomes

$$e^{jkz_0 y} \approx e^{jkz_0} e^{-j\frac{kz_0}{2} y^2} = e^{jkz_0} e^{-j\frac{kz_0}{2} (\xi\lambda)^2}$$

For a given weak phase grating,

$$U_{z=0}(x_1) = e^{jm \cos(2\pi x_1/T)} \approx 1 + jm \cos(2\pi x_1/T)$$

The angular spectrum at  $z=0$  is

$$A_{z=0}(\xi, \eta) = \iint_{-\infty}^{\infty} [1 + jm \cos(2\pi x_1/T)] e^{-j2\pi(\xi x_1 + \eta y_1)} dx_1 dy_1$$

$$= \left\{ \delta(\xi) + \frac{1}{2} jm \left[ \delta(\xi - \xi_0) + \delta(\xi + \xi_0) \right] \right\} \delta(\eta)$$

The angular spectrum at the observation plane is

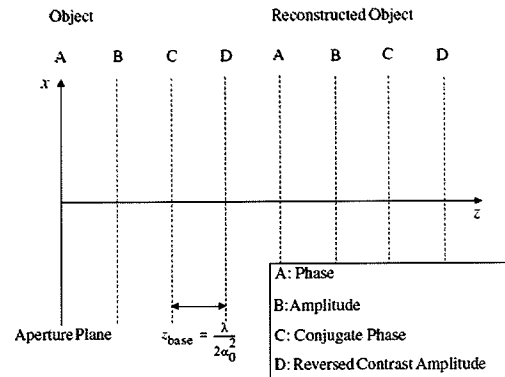
$$A(\xi, \eta; z_0) = e^{jkz_0} \left\{ \delta(\xi) + \frac{1}{2} jm e^{-j\frac{kz_0}{2} (\lambda\xi_0)^2} \left[ \delta(\xi - \xi_0) + \delta(\xi + \xi_0) \right] \right\} \delta(\eta)$$

$$= e^{jkz_0} \left\{ \delta(\xi) + \frac{1}{2} jm e^{j\psi} \left[ \delta(\xi - \xi_0) + \delta(\xi + \xi_0) \right] \right\} \delta(\eta)$$

It really depends on

$$\psi = -(kz_0/2)(\lambda\xi)^2$$

and the propagated field is repeated as



$$z_{\text{base}} = \frac{1}{2\lambda\xi_0^2} = \frac{\lambda}{2\alpha_0^2}$$

,where

#### 3.2 Babinet's Principle

Since diffraction is a linear and shift invariant process,

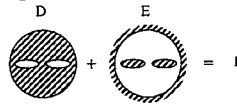
$$L(U_{\text{total}}) = L(U_1) + L(U_2) + \dots + L(U_n)$$

,where L: scalar diffraction linear operation

$U_{\text{total}}$ : total field transmitted through aperture

$U_1, U_2, \dots, U_n$ : Divided sub apertures

Then, the aperture algebra is very useful.

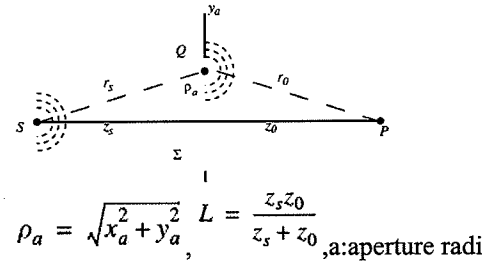


example:

Also, D and E is complementary apertures in this case.

#### 3.3 Fresnel Zones

Given aperture and source and observation positions as below



$$\rho_a = \sqrt{x_a^2 + y_a^2}, L = \frac{z_s z_o}{z_s + z_o}, a: \text{aperture radi}$$

The Fresnel number is

$$N_f = \frac{a^2}{\lambda L}$$

The radius of the  $m^{\text{th}}$  Fresnel zone is given by

$$\rho_m = \sqrt{m\lambda L}$$

the area of each Fresnel zone is

$$A_f = \pi\rho_{m+1}^2 - \pi\rho_m^2 = \pi\lambda L$$

#### 3.4 Diffraction behind circular aperture

Using the Fresnel zones, the axial field behind the circular aperture is

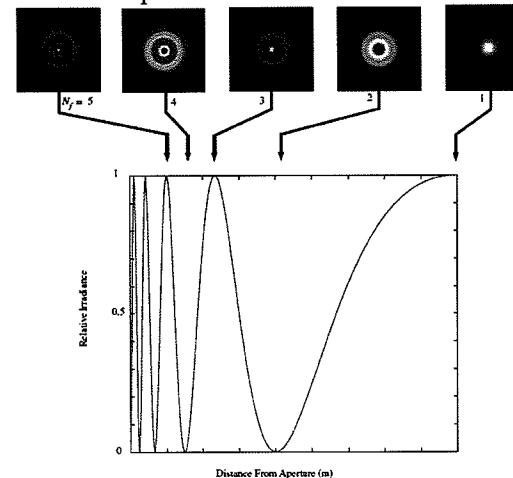
$$U(P) = -U_{\infty}(P)(e^{j\pi N_f} - 1)$$

,where  $U_{\infty}(P)$ : field w/o the aperture

Then, the axial irradiance becomes

$$I(P) = U(P)U^*(P) = 4I_{\infty}(P)\sin^2\left(\frac{\pi N_f}{2}\right)$$

Case 1) 1mm diameter aperture, illuminated with 0.5um plane wave for  $N_f = 1$  through 5.



case 2) converging wave illumination

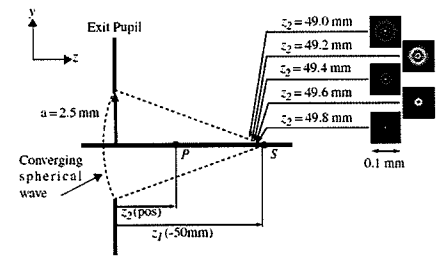
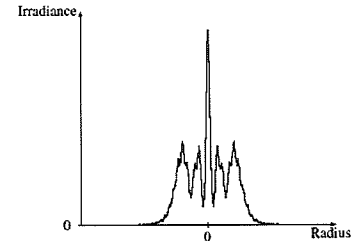


Fig. 6.36. Geometry for a laser beam converging through an aperture.  $\lambda = 0.5\mu\text{m}$ ,  $f = 50\text{mm}$  and  $a = 2.5\text{mm}$ .

The transverse irradiance also can be calculated. It has  $N_f (=5)$  peaks as below.



#### 3.5 Poisson's Spot

Behind a circular disk, on-axis field is

$$U_{\text{disk}}(P) = U_{\infty}(P) - U_{\text{aperture}}(P) = U_{\infty}(P) - U_{\infty}(P)(1 - e^{j\pi N_f}) = U_{\infty}(P)e^{j\pi N_f}$$

Thus,

$$I_{\text{disk}}(P) = I_{\infty}(P)$$

#### 3.6 Fresnel Zone Plate

If we put mask on even or odd Fresnel zones, it become the binary amplitude Fresnel Zone Plate. The on-axis irradiance becomes

$$I(P) = 4N_{FZ}^2 I_{\infty}(P)$$

,where  $N_{FZ}$ : number of open Fresnel zones

The transverse irradiance also can be calculated. It is has approximate width of the central spot

$$w = \frac{\lambda z_0}{a}$$

, which is close to the ideal lens case

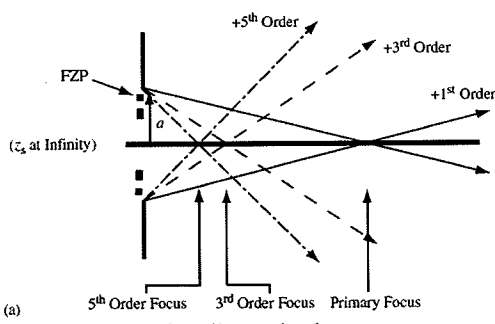
$$w = 1.22 \frac{\lambda z_0}{a}$$

, where w: width between zeros of the central spot in Airy disk

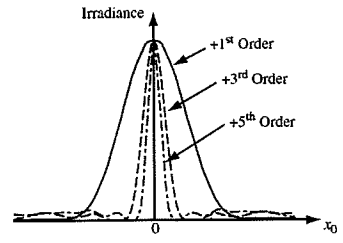
This Fresnel zone plate is very close to a lens, which has multiple focal lengths as

$$f_n = \frac{a^2 Z P 1}{n\lambda}$$

,where  $n = \dots, -5, -3, -1, 1, 3, 5 \dots$

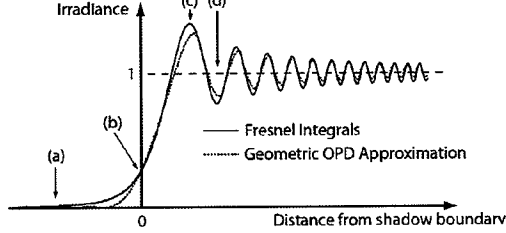


The transverse irradiance is shown as



### 3.7 Edge diffraction

The diffraction from an edge has irradiance



### 3.8 Fraunhofer diffraction

Case 1) Circular aperture w/ normal incident plane wave

$$U(\rho_0) = -\frac{je^{jkz_0}}{\lambda z_0} \exp\left(j \frac{k\rho_0^2}{2z_0}\right) \mathbf{B}_{\alpha_0=\frac{\rho_0}{\lambda z_0}} \left[ \text{circ}\left(\frac{\rho_1}{2a}\right) \right]$$

$$= -\frac{je^{jkz_0}}{\lambda z_0} \exp\left(j \frac{k\rho_0^2}{2z_0}\right) \pi a^2 \text{somb}\left(\frac{2a\rho_0}{\lambda z_0}\right)$$

Then, the irradiance becomes

$$I(\rho_0) = \left(\frac{1}{\lambda z_0}\right)^2 (\pi a^2)^2 \text{somb}^2\left(\frac{2a\rho_0}{\lambda z_0}\right)$$

We can also specify minimum resolvable distance

1) by Rayleigh's criterion

$$\rho_{RES} = 1.22 \frac{\lambda z_0}{2a} = 0.61 \frac{\lambda}{NA} = 1.22 \lambda f \#$$

2) by Sparrow criterion

$$\rho_{RES-SPARROW} = 0.94 \frac{\lambda z_0}{2a} = 0.47 \frac{\lambda}{NA} = 0.94 \lambda f \#$$

Case 2) Rectangular aperture w/ normal incident plane wave

$$U(x_0, y_0) = -\frac{je^{jkz_0}}{\lambda z_0} \exp\left[\frac{k}{2z_0}(x_0^2 + y_0^2)\right] \mathbf{F}_{\eta=\frac{y_0}{\lambda z_0}} \mathbf{F}_{\xi=\frac{x_0}{\lambda z_0}} \left[ A \text{rect}\left(\frac{x_1}{2a}\right) \text{rect}\left(\frac{y_1}{2b}\right) \right]$$

$$= -\frac{jAe^{jkz_0}}{\lambda z_0} \exp\left[\frac{k}{2z_0}(x_0^2 + y_0^2)\right] 4ab \text{sinc}\left(\frac{4ax_0}{\lambda z_0}\right) \text{sinc}\left(\frac{4by_0}{\lambda z_0}\right)$$

Then, the irradiance becomes

$$I(x_0, y_0) = \left(\frac{A}{\lambda z_0}\right)^2 (4ab)^2 \text{sinc}^2\left(\frac{2ax_0}{\lambda z_0}\right) \text{sinc}^2\left(\frac{2by_0}{\lambda z_0}\right)$$

The width between zeros of the central spot

$$s_x = \frac{\lambda z_0}{a} = \frac{\lambda}{NA_x}, \text{ and } s_y = \frac{\lambda z_0}{b} = \frac{\lambda}{NA_y}$$

Case 3) Double slit w/ normal incident plane wave

$$U(0, y_0) = -\frac{je^{jkz_0}}{\lambda z_0} \exp\left(\frac{ky_0^2}{2z_0}\right) \mathbf{F}_{\eta=\frac{y_0}{\lambda z_0}} \mathbf{F}_{\xi=\frac{x_0}{\lambda z_0}} \left\{ A \text{rect}\left(\frac{x_1}{2a}\right) \text{rect}\left(\frac{y_1}{2b}\right) \right\}$$

$$* \left[ \delta(y_1 + d/2) + \delta(y_1 - d/2) \right] \Big|_{x_0=0}$$

$$= -\frac{jAe^{jkz_0}}{\lambda z_0} \exp\left(\frac{ky_0^2}{2z_0}\right) 4ab \text{sinc}\left(\frac{2by_0}{\lambda z_0}\right) 2 \cos\left(\frac{2\pi dy_0}{2\lambda z_0}\right)$$

Then, the irradiance becomes

$$I(0, y_0) = 4 \left(\frac{A}{\lambda z_0}\right)^2 (4ab)^2 \text{sinc}^2\left(\frac{2by_0}{\lambda z_0}\right) \cos^2\left(\frac{2\pi dy_0}{2\lambda z_0}\right)$$

$$= 4 \left(\frac{A}{\lambda z_0}\right)^2 (4ab)^2 \text{sinc}^2\left(\frac{2by_0}{\lambda z_0}\right) \left[ \frac{1}{2} + \frac{1}{2} \cos\left(\frac{2\pi dy_0}{\lambda z_0}\right) \right]$$

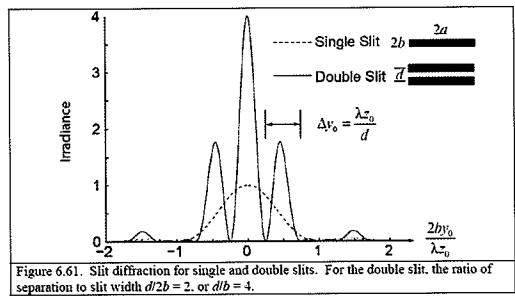


Figure 6.61. Slit diffraction for single and double slits. For the double slit, the ratio of separation to slit width  $d/2b = 2$ , or  $d/b = 4$ .

Case 4) N slits (or grating)

For N slits, the number of secondary peaks between the primary peaks equals N-2

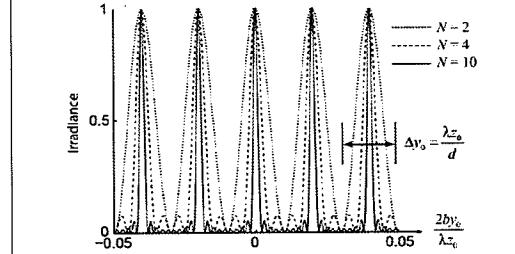


Figure 6.63. Increasing the number of slits N. The number of secondary peaks between the primary peaks equals N-2. In this figure, the peaks are normalized to unity irradiance. In reality, the peak irradiance increases as N^2. Also,  $d/b = 100$ .

For gratings, it follows the grating equation

$$n_s \sin \theta_s + n_0 \sin \theta_m = \frac{m\lambda}{d}$$

For a given transmission of aperture

$$t(x_1, y_1) = \left[ f_d(y_1) * \frac{1}{d} \text{comb}\left(\frac{y_1}{d}\right) \right] \frac{1}{2a} \text{circ}\left(\frac{\sqrt{x_1^2 + y_1^2}}{2a}\right)$$

and given plane wave illumination

$$U(x_1, y_1) = A \exp\left(j2\pi \frac{\beta_s}{\lambda} y_1\right)$$

The irradiances is (for  $2a \gg d$ )

$$I(x_0, y_0) \approx \left(\frac{\pi a^2 A}{\lambda z_0 d}\right)^2 \sum_{m=-\infty}^{\infty} \left| \mathbf{F}_m[f_d(y_1)] \right|^2 \text{somb}^2\left[\frac{2a}{\lambda z_0} \sqrt{x_0^2 + \left(y_0 + \frac{m\lambda z_0}{d} - \beta_s z_0\right)^2}\right]$$

and the absolute diffraction efficiency is

$$E_m = \left| \frac{1}{d} \mathbf{F}_m[f_d(y_1)] \right|^2$$

For example, consider a sine wave grating with

$$f_d(y_1) = \left[ \frac{1}{2} + \frac{1}{2} \sin(2\pi y_1/d) \right] \text{rect}\left(\frac{y_1}{d}\right)$$

so

$$E_m = \left| \frac{1}{d} \mathbf{F}_m \left\{ \left[ \frac{1}{2} + \frac{1}{2} \sin(2\pi y_1/d) \right] \text{rect}\left(\frac{y_1}{d}\right) \right\} \right|^2$$

$$= \left| \frac{1}{2} \text{sinc}(m) + \frac{\text{sinc}(m-1) - \text{sinc}(m+1)}{4j} \right|^2$$

$$E_0 = 0.25 \text{ (25\%)} \text{ and } E_{\pm 1} = 0.0625 \text{ (6.25\%)}$$

case 5) Wide-angle diffraction to a spherical surface

Gratings often diffract light over large angles, which are not strictly allowed in the Fraunhofer development in previous section. This restrictions are removed by developing a diffraction formula for propagating from a plane to a spherical surface.

For a observation surface, which is a sphere of radius r, and centered at  $(x_1, y_1) = (0, 0)$ , the field is

$$U(\xi, \eta) = -\frac{j\gamma e^{jkr}}{\lambda r} \mathbf{F}_{\beta=\eta} \mathbf{F}_{\alpha=\xi} \left[ U_{z=0}(x_1, y_1) \right]$$

The observation coordinate system is defined on a hemisphere

$$x_0 = \lambda \xi r$$

$$y_0 = \lambda \eta r$$

$$z_0 = \gamma r$$

, where

case 6) Conical diffraction

Illumination plane wave propagation vector is not in the plane defined by g (grating vector) and the normal to the grating surface.

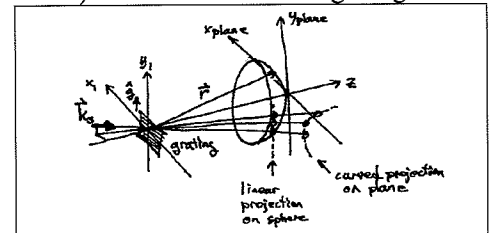


Figure 6.72. Conical diffraction. Projection of diffraction orders onto the observation hemisphere and the flat vertex plane.

The conical nature is a result of observing the diffraction on a plane, rather than a sphere. The diffraction orders are said to be linear in "direction cosine space."